

NEW YORK ACADEMY OF SCIENCES.

A meeting of the New York Academy of Sciences, formerly the Lyceum of Natural History, was held recently in Professor Mayer's lecture room, at the Stevens Institute of Technology. Two communications were made by Professor A. M. Mayer and one by President Henry Morton. The first paper was by Professor Mayer on

CROOKES' RADIOMETER,

and on some results obtained by the action of sound pulses on an apparatus constructed in a similar manner. Crookes was led to the construction of his delicate instrument in the following manner: He made a torsion balance consisting of a bar of pith suspended by a fine filament. One half the length of this bar was blackened. On exposing this apparatus to rays from different parts of the prismatic spectrum, he found that the torsion balance moved through spaces proportional to the thermometric effect produced by the same area of rays falling on a thermopile. Representing the motion produced by the ultra-red rays by 100, those of the other rays were as follows:

Extreme red.....	85	Blue.....	22
Red.....	73	Indigo.....	8½
Orange.....	66	Violet.....	6
Yellow.....	57	Ultra violet.....	5
Green.....	41		

The difficulty of ascertaining facts in Nature, even by the most careful observers, is well illustrated here by the fact that Crookes overlooked the circumstance that he was here operating with a purely accidental spectrum, the proportion of whose parts depended entirely on the nature of his prism. Had he employed the normal or diffraction spectrum produced by the passage of light through finely ruled glass plates, his results would have been the same as those obtained many years ago in this country by Dr. Draper.

Crookes found that the bar of the torsion balance was attracted when the apparatus contained air, and repelled when it was placed in a vacuum; also that the radiation from a candle on blackened pith was 5½ times what it was on plain pith. From these observations to the construction of the radiometer was but a short step. Two fine wires, at right angles to each other, were provided with little vanes of mica blackened on one side, and the whole suspended on a pivot and enclosed in a glass vessel, from which the air was exhausted by means of a Sprengel pump. When rays of light fall on this apparatus, the differential action of the blackened and natural surfaces of the vanes gives rise to a continuous rotation. The rate of this rotation, and hence the intensity of the exciting cause, was obtained by means of a small electromagnet placed in the apparatus in such a manner as to register the number of revolutions by making a series of dots on a slip of paper.

Professor Mayer then exhibited this very delicate instrument to the meeting, and showed the increase in the velocity of rotation on bringing it nearer a source of light.

In order to measure the repulsion of a blackened surface in a vacuum, Crookes employed W. Ritchie's torsion balance, described in the "Transactions of the Royal Society" of 1830. This is so arranged that the repulsion of the blackened surface twists a fine glass thread, to which is attached a mirror projecting a beam of light on a screen. By means of a screw, the circumference of which is divided into 360°, the glass thread is turned back again until the beam occupies its original position. The amount of torsion is then read off on the screw. The extreme delicacy of this instrument may be appreciated from the fact that the $\frac{1}{100}$ of a grain produces a torsion through 10,000 degrees or about 28 rotations of the thread. As it is sensitive to 1° of rotation, it is evident that we can thus weigh $\frac{1}{100000}$ of $\frac{1}{100} = \frac{1}{10000000}$ of a grain!

A candle at a distance of 6 inches repels 2 square inches of surface of blackened pith with a force of 0.001772 grain, at 12 inches distance with a force of 0.000444 grain. Starting out with the latter figure, and remembering that the effects are as the reciprocals of the squares of the distances, we should obtain at 6 inches a force of 0.001776: which differs from the result actually obtained by experiment by only the $\frac{4}{1000000}$ of a grain.

After succeeding in constructing so perfect an instrument, it is not surprising that Crookes should be elated; but it is to be regretted that he should express himself as he did in the following extract from his paper: "A candle 12 inches off, acting on 2 square inches of surface, was found equal to 0.000444 grains; the sun, equalling 1,000 candles at 12 inches, gives a pressure of 0.44000 grain; that is equal to about 32 grains per square foot, to 2 cwt. per acre, to 57 tons per square mile, or nearly 3,000,000,000 tons on the exposed surface of the globe—sufficient to knock the earth out of its orbit if it came upon it suddenly." It is true he immediately modifies this statement, but it is liable to be quoted without the following disclaimer: "It must be remembered that our earth is not a lamp-blackened body enclosed in a glass case, nor is its shape such as to give the maximum surface with the minimum of weight."

Still more mischievous, however, is the pretention to "weigh a beam of light." That it is not light which causes the rotation in Crookes' radiometer has been conclusively proved by Schuster ("Proceedings of the Royal Society," April, 1876). He made a radiometer having one arm a magnet. When a strong light fell on this apparatus, it overcame the directive force of the magnet, and caused it to rotate in the usual way. Now on floating the instrument in water, and holding a magnet outside to keep it stationary in spite of the strong light falling upon it, the bulb began to rotate in a direction opposite to that which the light would have imparted to the vanes.

The only plausible explanation hitherto offered of the instrument is the following: The vacuum of the Sprengel pump is not a perfect one; but in the highly rarefied air contained in the bulb, the molecules have a much greater amplitude of swing than in their ordinary condition, there being vastly fewer in the same space. Hence the currents set up by the very feeble heat of the rays of light are sufficient to produce a much more intense action than could take place in dense air. The blackened surfaces become heated more than the natural ones, and consequently repel more particles, the reaction of which causes motion; and the apparatus being free to turn, the motion becomes one of rotation, which continues until the effects on the two surfaces become equalized.

To show that the effect was due to heat, Professor Mayer interposed a glass plate between the apparatus and the diffused light from a window. The effect was to stop the rotation by cutting off the heat rays, while the light rays passed through freely. Rotation was also produced by placing the hand near the radiometer.

AN ANALOGOUS APPARATUS FOR SOUND.

In the next place, Professor Mayer exhibited an apparatus constructed by him to produce motion by means of sound pulses. Four glass resonators on cross arms were suspended by means of a string. On sounding an organ pipe in tune with the resonators, and bringing it opposite the mouth of one of them, the resonator was repelled and the apparatus commenced to rotate. This experiment was the more striking from the fact that, so far from any current of air proceeding out of the mouth of the organ pipe, the air is actually sucked in, as may be rendered visible by means of smoke from a cigar. The smoke is carried up the pipe even when the latter is closed at the top with cotton wool so as to smother the sound. On substituting disks of cardboard for the resonators, they were drawn up to the mouth of the organ pipe with considerable force. When fine silica powder was placed in the resonators, it was thrown into violent motion on sounding the pipe.

OBLITERATION OF SONOROUS SENSATIONS.

Professor Mayer next described some interesting experiments, the effect of which will be to modify our present theories of audition. He took an American clock ticking 4 beats a second and a watch ticking 5. At a distance of several feet from the ear, all the ticks of both could be plainly heard, even those which coincided. On gradually moving the watch away from the ear, a point was reached where the fifth or coincident tick of the watch became inaudible. A watch beating 4 times a second was then substituted and set to gain 30 seconds an hour on the clock. Coincidence then occurred every 2 minutes. Removing the watch 24 inches from the ear, its ticks became extinguished for 9 seconds at a time.

To determine the relative intensities of these ticks, numerous experiments were made on still nights in the country, and it was found that the ticks of the clock became inaudible at a distance of 350 feet, and those of the watch at a distance of 20 feet. The ratio of the squares of these numbers makes the ticks of the clock about 300 times more intense than those of the watch. Standing at different distances from the clock, and holding a slender stick graduated to inches and tenths against his zygomatic process, the experimenter slid a watch along the stick (taking care not to touch it) until its fifth tick disappeared. The relative intensities of clock and watch ticks at the same distance being known, their intensities at the distances in these experiments were obtained by the law of reciprocal squares. The slightest noise or breeze renders the experiment impossible.

In the same way one musical sound may obliterate the sensation of one higher than itself. To obviate the objection that this might be due to a change of timbre, Professor Mayer caused one of the notes to sound periodically by opening and shutting the resonant box of a fork with the hand at regular intervals. Now on sounding another lower note, either from a fork or an organ pipe, continuously, its timbre could only be changed during the sounding intervals of the other note, and would be restored during the intervals of silence, if the objection is just. No such effect, however, is observed. The actual result is that the sound of the fork becomes more and more feeble, and finally its sensation is entirely lost. If at this point the other sound is stopped, it is found that the fork is still vibrating and emitting a sound, which had been completely overpowered by the graver one of the pipe. The same takes place if a closed organ pipe of the same pitch is substituted for the fork and its sound is made intermittent.

Another very remarkable fact is that even a very intense sound cannot obliterate the sensation of one lower in pitch. When the ear ceases to perceive the lower sound, it is generally found, on stopping the higher one, that the lower fork has entirely ceased to vibrate. Numerous experiments have been made, through a range of four octaves and on a score of ears, always with the same result.

If a sentence is read over and over in the same tone of voice and modulation, while an Ut_2 pipe is sounding, it seems as if two persons were reading, one with a very grave voice (consisting of all the vocal sounds below Ut_2) and one with a high, squeaky, and nasal voice. The intermediate harmonies are obliterated. The fundamental tone of most notes is perceived as the strongest, because each harmonic is diminished by the sum of all below it, while the lowest is not affected by any of them. On sounding Ut_2 together with an Ut_2 free reed pipe, the component Ut_2 is perceived unaffected in intensity, while all the higher harmonics are obliterated except Mi_2 and Sol_2 , which are very distinct. The deep residual sound of a New York fire alarm bell is heard

distinctly through all the noises of the street higher than it is until it dies entirely away. It seems in these experiments as though we had a special sense for the lower sounds.

Professor Mayer stated that since these discoveries he has repeatedly witnessed in orchestral music the entire obliteration of all sounds from the violins by the deeper and more intense sounds of the wind instruments, while the bass viols hold their own. From the same cause the clarinets lose their peculiar quality and charm. The leader probably heard them all; but as the ears of the audience hold different relation of distance to the several instruments, they are actually paying for his enjoyment instead of their own. His position ought therefore to be that of an average hearer. The laws of orchestration can only be arrived at by a quantitative determination of the intensities of all musical instruments. When quantitative analyses of the compound sounds in the different instruments shall have been made, the composer will be better able to command such qualities of sound as he desires to bring out. It is evident that instruments to be used in orchestras must be differently constructed from those designed for solos or quartets, and that they must have their peculiarities of timbre exaggerated.

Professor Mayer concluded by stating that he is at present engaged in determining the intensity of sounds either existing alone or as components of compound notes.

The meeting then adjourned to President Morton's lecture room to listen to a communication on some fluorescent bodies.

FLUORESCENCE.

Professor Morton began by stating that the property of lengthening the wave length of light rays, to which fluorescence was due, was one inherent in a great range of bodies. Fluorescence is produced when light rays of one kind are converted into another having a longer wave; when the rays are lengthened so much as to exceed the limits of the spectrum, they are converted into heat. A number of fluorescing substances were then exhibited, producing effects of indescribable beauty.

Fluoresceine, a substance having little color of its own, fluoresced with a rich bright green when illuminated with the electric light passing through blue or violet glass.

Eocene, the tetrabromide of fluoresceine, which has naturally a delicate pink color in solution, became dark green under the same conditions. In this, as in the other solutions shown, the fluorescence produced a very noticeable opacity.

Purpurine, one of the coloring substances made from madder, became very opaque orange yellow. The tinctorial power of this and some of the other substances shown is perfectly marvelous. As much as will lie on the point of a penknife, dissolved in a gallon of water, had to be diluted again with three or four times as much water to produce a good effect. Fluorescence enables us here to distinguish with perfect certainty between the natural product and artificial purpurine. A piece of a dressing gown dyed with madder, having passed the last stage of its usefulness, had been employed for a long time as a rag to wipe up things in the laboratory. A bit of this was treated to extract the purpurine, and its fluorescence was shown to be perfectly distinct and characteristic. By thus analyzing a garment dyed with artificial preparations, we would be enabled to decide with certainty that it could not have been manufactured before the date of the invention of the artificial dyestuffs.

Chlorophyll, made by steeping exhausted tea leaves in alcohol, is a liquid of a very deep green, so deep indeed as to appear almost black. It fluoresced with a magnificent crimson.

A solution of bisulpho-bichlor-anthracenic acid, which is perfectly colorless, fluoresced with a purplish blue.

A dish of nitrate of uranium crystals fluoresced with a luminous yellow color.

An extremely beautiful effect was produced by illuminating a screen on which a large wreath encircling the monogram of the Stevens Institute of Technology was painted by means of different fluorescing substances.

President Morton concluded by exhibiting some absorption bands, and pointing out how they serve to distinguish the substances by which they are produced.

ASTRONOMICAL NOTES.

OBSERVATORY OF VASSAR COLLEGE.

The computations and some of the observations in the following notes are from students in the astronomical department. The times of risings and settings of planets are approximate, but sufficiently accurate to enable an ordinary observer to find the object mentioned. M. M.

Positions of Planets for July, 1876.

Mercury.

On July 1, Mercury rises at 3h. 33m. A. M., and sets at 5h. 56m. P. M. On July 31, Mercury rises at 4h. 29m. A. M., and sets at 7h. 6m. P. M.

Mercury is at its greatest elongation west of the sun on the 8th, and can then be best seen. It should be looked for before sunrise, a little south of the point of sunrise.

Venus.

On July 1 Venus rises at 6h. 13m. A. M., and sets at 8h. 30m. P. M. On the 31st, Venus rises at 3h. 21m. A. M., and sets at 5h. 20m. P. M.

Venus is becoming less and less conspicuous, and will probably not be noticed after the first week of July, being too nearly in range with the sun. After the middle of the month Venus should be looked for in the morning, as is has passed to the west of the sun.

Mars.

Mars is now very small, and will not be noticed by ordinary observers. It will be very near Venus on the 5th, and may possibly be seen just after sunset some 45° farther

